

PSII-63 AFES (GCM) LETKF Data Assimilation System for Venus Norihiko Sugimoto (Keio Univ.),



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SOUNDER (DAY & NIGHT

100m/s !!

60 times faster

than surface

rotation rate

100

75

V9 & V10

1. Introduction

Venus atmosphere: Fast zonal flow (Super-rotation; SR)



	Venus	Earth
Radius	6050 km	6378 km
Revolution	224 days	365 days
Rotation	243 days (1.8m/s)	1 day (460m/s)
1 solar day	117 days	1 day
Composition	CO ₂	N ₂ , O ₂
Albedo	0.78	0.3
Surface pres.	92 bar	1 bar



east

Mechanism of the super rotation (reproduced in VGCM)

1. Mean meridional circulation 2. Thermal tide (Fels & Lindzen, 1986)



3. ALEDAS-V (AFES LETKF Data Assimilation System for Venus)

Data assimilation: First analysis for the Venus atmosphere



http://www.data-assimilation.riken.jp/jp/research/research.html

Mars Atmosphere Data Assimilation workshop (2018)

Akatsuki (Venus climate orbiter): Frequent observations at multiple altitudes







2. AFES-Venus (Atmospheric GCM For the Earth Simulator for Venus)

- Model description
 - ✓ 3-D Primitive equation on sphere (hydro static balance) without moist processes
 - ✓ Resolution: T42L60, T159L120 to T639L260 ($\Delta x \sim 20 \text{km} \Delta z \sim 0.25 \text{km}$)...
 - ✓ Specific heat: Cp is constant (1000 Jkg⁻¹k⁻¹)
 - ✓ Horizontal hyper-viscosity: 0.1 (T42) to 0.001 (T639) Earth days for 1/e
 - ✓ Vertical eddy viscosity: $0.15 0.0015 \text{ m}^2\text{s}^{-1}$; 10 values
 - ✓ Rayleigh friction: lowest and above 80 km (sponge layer except for zonal flow)
 - ✓ No topography and planetary boundary layer + cloud physics

Solar heating

- ✓ Zonal and/or diurnal component of realistic heating (Tomasko et al., 1980; Crisp, 1986)
- Infrared radiative process
 - ✓ Simplified by Newtonian cooling: $dT/dt = -\kappa (T-Tref(z))$, κ : based on Crisp (1986)
 - \checkmark Tref(z): horizontally uniform field but realistic vertical profile of static stability





- Local: considers only observations within a certain distance. Ensemble: uses an ensemble of GCM forecasts. Transform: uses a square-root filter. Kalman Filter: uses past information to update the present state, and estimates both the state and its uncertainty (covariance)
- Data assimilation for real observations
 - ✓ Horizontal winds from Venus Express (Sugimoto et al., GRL2019b)

Time

- Horizontal winds from Akatsuki (Fujisawa et al., Sci. Rep.2022) \checkmark
- ✓ Cold collar in the first analysis (*Ando et al., JGR2023*)



- Sugimoto et al. (SREP2017) observations observations (horizontal winds ✓ 9-hour forecast from t=0, input observations from t=3 to 9
 - ✓ Output analysis at t=6 (=4D LETKF)

Test of the horizontal winds assimilation (Sugimoto+2019b GRL) Cloud tracking of the ultraviolet images (UVI) by "Venus Express"



Initial condition

(1) Motionless state

- Mimic observed temperature field including cloud layer with low static stability: $\Gamma(z) = dT/dz + g/Cp$
- (2) Idealized SR (solid body rotation)
 - Zonal flow increases with height linearly from ground to 70 km. 100 m/s above 70 km (const.).
 - Temperature field is in balance with zonal flow field (gradient wind balance).

Strategy

- Start from idealized super rotation
 - ✓ Saving computational cost for high resolution run
- Maintain super rotation with realistic setting ✓ under the realistic solar heating and static stability without artificial f

Targets

- Focus on atmospheric motions near the cloud level
 - ✓ Baroclinic instability; Not observed but predicted theoretically (Sugimoto et al., JGR2014) Taylor et al. (1980)
 - ✓ Neutral waves; Observed by cloud images but unexplained (Sugimoto et al., GRL2014)
 - ✓ Thermal tide; Elucidate its role, horizontal and vertical structures (*Takagi et al., JGR2018*)
 - ✓ Energy spectra; Traditional analysis on Earth but no Venus case (*Kashimura et al., in prep.*)
 - ✓ Polar vortex; "Axi-asymmetric" structure observed in VIRTIS (Ando et al., JGR2017)
 - ✓ Cold collar; Cold latitudinal band not reproduced in GCMs (Ando et al., Nature Com.2016)
 - ✓ Planetary scale streak structure... (Kashimura et al., Nature Com. 2019)







225 2

240 2 2

250 259

r(⁰κ) 230 23 235 240

Ando et al. (Nature Com. 2016)



Belton et al. (1991)

(a) U, V, Z at 70 km, T at 60 km, day 47

Zonal wind: latitude-altitude cross section Comparison with AFES-Venus (without DA) • AFES-Venus: The zonal wind is about 130 m/s at about 70 km. · Analysis: The zonal wind is globally reduced Comparison with observation • Observations: The zonal wind is 80-100 m/s at about 70 km. Too fast super-rotation in AFES-Venus is improved globally.



- Observing System Simulation Experiments (OSSEs)
 - ✓ Cross-link Radio Occultation measurements (Fujisawa et al., Icarus, 2023)
 - Ultra-Violet Images for Kelvins waves (Sugimoto et al., Atmosphere2021, 2022) \checkmark
 - ✓ Ultra-Violet Images for Rossby waves (*Komori et al., in preparation*)
 - ✓ Long Infrared images for thermal tides (*Sugimoto et al., GeoSci. Lett., 2022*)





First analysis for the Venus atmosphere (Press released in Sep., 2022)

@https://www.jamstec.go.jp/es/en/







Akatsuki Infrared observations

Kashimura et al. (Nature Com. 2019)

Recent progress

- ✓ Fully developed super-rotation driven by MMC (*Sugimoto et al., GRL2019a*)
- ✓ Super-rotation independent of horizontal diffusion (Sugimoto et al., EPS2023)
- ✓ Static stability compared with observations (Ando et al., Sci. Rep. 2020, JGR 2022)
- ✓ Generation of gravity waves from thermal tides (*Sugimoto et al., N.Com., 2021*)
- ✓ Implementation of cloud physics and its variation (Ando et al., JGR2020, 2021)
- ✓ Improvement of static stability and thermal tides (Suzuki et al., JGR2022)
- ✓ Reproduction of 4-Day and 5-Day Waves (*Takagi et al., JGR2022*)







-0.5

0.5

assimilated several real observations and conducted OSSEs for future missions.

It would be possible to assimilate other obs. directly, such as cloud opacity and radiance, by improvement of physical processes of AVES-Venus...